

PHYSICS 526 NOTES: FLUID DYNAMICS

Jean Eilek

Physics Department, New Mexico Tech
Socorro, NM 87801, U.S.A. and
jeilek@aoc.nrao.edu

Our goal in this course is to explore the basic physics of fluid systems, both with and without the effects of magnetic fields. Fluid physics by itself – hydrodynamics – applies to any neutral fluid (think about water, or molasses, or the earth’s atmosphere), whether or not there are magnetic fields around. “Hydro” or “HD” has broad applications – smooth flows, turbulent flows, shocks, sound waves, instabilities. If the fluid is ionized (think of liquid sodium, the higher reaches of the earth’s atmosphere, or just about any astrophysical system you can bring to mind), it can (and almost certainly will) carry a current. The interaction of the current with a magnetic field (external or self-generated) modifies all of the phenomena above and adds some new ones. This is described by magnetohydrodynamics (MHD).

Traditionally, HD and MHD are treated separately, but they have a lot in common, and I have learned a lot by comparing and contrasting HD and MHD phenomena. I’m hoping to share that with you as we go through this course.

A word of caution: you should especially note *units and dimensions*. These notes are in cgs. That makes very little difference for “rocks” (analyses that involve mass, length, time); but it makes a big difference for electrodynamics and MHD. The \mathbf{E} and \mathbf{B} fields, as well as the fundamental charge, have different dimensions in cgs than in SI; and the coupling constants in Maxwell’s equations are different. Both systems appear in the literature; while there is some trend towards SI, many important references are in cgs. I am going to follow my experience and preference and use cgs; I’ll also give the SI versions of most critical equations.

Contents

1. BASIC CONCEPTS AND TOOLS	1
A. Kinematics: How to Describe a Flow	1
stream function	1
B. Mass Conservation: The Continuity Equation	1
lagrangian derivative	2
C. Momentum Conservation: Euler’s Equation	2
1. control volumes	3
2. bernoulli’s relation	3
D. Work in a Rotating Frame	3
E. Dimensional Analysis	4
F. Appendix: When can we use hydrodynamics?	5
1. hard sphere collisions	5
2. plasmas: the coulomb cross section	6
3. collisionless plasmas	7
2. VISCOSITY AND LAMINAR FLOW: BASICS	8
A. One-dimensional Laminar Flows	8
B. Steady Flow: Cartesian Applications	9
1. flow between parallel plates	9
2. flow in an open channel	9
3. hele shaw flow	9
C. Steady Flow: Cylindrical Applications	10
1. pipe flow	10
2. circulating flow	10
D. Viscous Stresses, Generally	11
1. do it physically first	11
2. then do it formally	12
E. The Navier-Stokes Equation (in Cartesian)	13
F. Appendix: Navier-Stokes in other coordinates	13
1. Cylindrical coordinates	13
2. Spherical Polar Coordinates	14
3. POTENTIAL FLOW	16
A. Setups and Assumptions	16
1. velocity potential	16
2. stream function	16
3. other coordinates	16
B. Two-dimensional Planar Problems	17
1. sources and sinks, flow past half body	17
2. flow past a cylinder	18
C. Axisymmetric 3D Problems	18
point source and stream flow	19
flow around a sphere	19
a line source	19
D. d’Alembert’s “Paradox”	20
4. VORTICITY	21
A. Vortex Kinematics	21
B. Vortex Dynamics	21
C. Conservation of Circulation: Kelvin’s Theorem	22
D. The Magnus Effect	23
E. Vortex lines and their behavior	24

F. Generation of Vorticity	25	strong shock limit, normal shocks	49
5. LAMINAR FLOW: MORE APPLICATIONS	26	C. Oblique shocks	49
A. Geostrophic Flow	26	1. two possible deflections	50
atmospheric circulation	26	2. high mach number limit	51
rossby waves	26	D. The Weak Shock Limit	51
the ekman layer	27	prandtl-meyer function	51
B. Viscous Flow: Time-Dependent Problems	27	10. ONE-DIMENSIONAL UNSTEADY FLOW	53
1. similarity methods in a diffusion equation	28	A. Shock Formation: The Physical Picture	53
2. smoothing out a velocity jump	28	shock thickness	53
3. flow above an oscillating plate	28	B. The Method of Characteristics	54
4. irrotational vortex decay	28	1. piston motion in a channel	55
C. Creeping Flow	29	2. connection to shock formation	56
6. BASICS OF COMPRESSIBLE FLOW	31	3. traffic shocks	57
A. Some useful thermodynamic quantities	31	11. TWO-DIMENSIONAL STEADY FLOW	58
B. Hydrostatics: gaseous atmospheres	31	A. The Nature of Steady, two-dimensional flow.	58
1. constant gravity: the exponential atmosphere	31	B. Signal Propagation in Flows	58
2. variable gravity: the isothermal sphere	31	C. How Does Supersonic Flow Turn a Corner?	59
3. reality: nonisothermal atmospheres	32	D. One example: Prandtl-Meyer flow	60
4. adiabatic atmosphere	32	12. SIMILARITY SOLUTIONS	62
C. Convective Stability	33	A. Blast Waves: the Sedov-Taylor Solution	62
1. adiabatic atmosphere	34	B. Prandtl-Meyer flow, revisited	64
2. potential temperature	34	C. Gravitational Collapse: the Shu Solution	65
3. brunt-väisälä frequency	34	13. FOUNDATIONS OF MHD	66
D. Energetics of Compressible Flow	35	A. Remember your E&M?	66
E. Appendix: Viscous Dissipation	36	1. conductivity and ohm's law: i	66
7. SIGNAL PROPAGATION	38	2. conductivity and ohm's law: ii	66
A. Sound Waves and the Signal Speed	38	B. Field Evolution: Induction Equation	66
sound waves: a physical approach	38	1. ideal limit: flux freezing	67
sound waves: a formal approach	38	2. resistive limit: flux annihilation	67
B. Why is the sound speed important?	39	C. Fluid Equations: Lorentz force	68
1. when can we assume incompressible flow?	39	D. Fluid Equations: Energetics	68
2. the importance of causality	39	1. energetics of the e and b fields	68
C. Weak Waves and Causality	40	2. energetics of the fluid	68
D. Two examples of simple waves	41	E. Appendix I: Conductivity notes	69
1. shock tube	41	1. collisional conductivity	69
2. piston problem	41	2. cross-field conductivity	70
3. waves at boundaries	41	F. Appendix II: Do It in SI	71
8. ONE-DIMENSIONAL STEADY FLOW	43	1. maxwell in SI:	71
A. Two-point Connections in Steady Flow	43	2. induction equation	71
B. One-dimensional channel flow	44	3. force equation, magnetic tension and pressure	71
1. nozzles and diffusers	44	4. energy equation	71
2. a smooth transition?	44	14. SIMPLE MHD EQUILIBRIA	72
3. normal shocks	45	A. Potential Fields	72
4. 1D flows with internal shocks	45	B. Plasma Confinement	72
C. Spherical stellar wind flow	46	1. theta pinch	72
9. SHOCKS IN FLUID FLOW	48	2. bennet pinch or z pinch	73
A. Jump conditions	48	3. general screw pinch	74
B. Normal shocks	48	C. Force-Free Fields	74

1. cylindrical geometry	74	A. Advection vs. dissipation of magnetic field	96
2. spherical geometry	74	1. resistive limit: field decay on axis	97
3. non-linear fields	75	2. advection limit: field growth on axis	97
4. boundaries	75	B. Steady, 2D Reconnection	97
D. Gravitational Equilibrium I	75	1. why do field lines “break”?	97
1. planar geometry: the galaxy	75	2. sweet-parker model	98
E. Magnetic Bouyancy	76	3. energetics	99
1. “convective” instability	76	4. timescales and rates	99
2. scale of unstable perturbations: Parker Instability	77	C. Speed up the reconnection?	99
15. NOT-SO-SIMPLE EQUILIBRIA	78	1. petschek reconnection	99
A. Flux functions I: Gravitational Equilibrium Revisited	78	2. compressible flow	100
1. define the flux functions	78	3. anomalous diffusion	100
2. apply them: solar magnetic arches	79	D. Other approaches	100
B. Flux Functions II: Grad-Shafranov Equation	79	1. spontaneous reconnection	100
1. define the flux functions	80	2. driven reconnection	101
2. apply them: the g-s equation	80	3. non-steady reconnection	101
3. example: simple pinches	81	4. three-dimensional reconnection	101
C. Helicity and Taylor relaxation	81	E. Appendix	101
1. helicity	81	1. diffusion-only solution	101
2. Invariance of the helicity	82	2. advection-only solution	102
3. the minimum energy state	82	19. FLUID INSTABILITIES	104
4. taylor relaxation	83	A. Buoyancy and Thermal Convection	104
16. MHD EFFECTS IN FLUID FLOWS	85	B. The Rayleigh Taylor Instability	104
A. Magnetic damping and stirring	85	1. the physics	104
1. Magnetic damping.	85	2. the math	105
2. Magnetic stirring.	85	3. the magnetized case	106
B. Channel flow with MHD: Hartmann flow	85	C. The Kelvin Helmholtz Instability	107
1. The basic setup	86	1. the physics	107
2. the solution	86	2. the math	107
3. an mhd generator or an mhd pump?	86	3. the magnetized case	108
C. Magnetic Coupling of Two Plasma Slabs.	87	20. IDEAL MHD INSTABILITIES	110
1. pure mhd	87	A. Overview; Energy Methods	110
2. current-based approach	87	1. energy methods	110
3. energetics and equivalent circuit	88	2. the details	110
17. WAVES AND SHOCKS IN MHD	89	B. Apply: Pinch Instabilities	111
A. MHD Waves	89	1. theta pinch	111
1. basic structure: linear analysis	89	2. z pinch	112
2. alfvén waves	89	21. RESISTIVE MHD INSTABILITIES	114
3. magnetosonic waves	90	A. Tearing Mode: the Physics	114
4. validity of mhd wave theory	91	B. Tearing Mode: the Math	115
B. MHD Shocks; Jump Conditions	91	C. Tearing Mode: the Consequences	116
C. Perpendicular (Normal) Shocks	92	22. TURBULENCE, part I	117
D. Oblique Shocks	93	A. The transition to turbulence	117
1. do it generally	93	B. Turbulent flows: overview	118
2. find a useful reference frame	94	1. characteristics	118
3. now solve the system: i	94	2. behavior	119
4. now solve the system: ii	95	C. Homogeneous Turbulence	119
5. back to the physics	95	1. overview	119
18. MAGNETIC RECONNECTION	96	2. eddies and the energy cascade	120

3. the kolmogorov scaling arguments	120
4. what if the fluid is magnetized?	121
D. Appendix: fun facts from Fourier transforms	121
23. TURBULENCE, part II	123
A. Mean Field Equations	123
1. the continuity equation	123
2. the mean momentum equation	123
3. example: 2D channel flow	124
4. the mean energy equation	124
5. what about the turbulent terms?	124
B. Two-dimensional Turbulence	125
C. Small scales and intermittency	126
1. intermittency	126
2. turbulence on small scales	126
3. the role of vortex filaments	126
24. MHD Turbulence and Dynamos	128
A. Magnetic fields in Isotropic turbulence	128
B. The Inertial Range in MHD Turbulence	128
1. the kraichnan model	129
2. anisotropy; goldreich-sridhar model	129
3. what now?	129
C. Cascades and Related Things	130
1. ideal invariants	130
2. cascade directions	130
3. self-organization in mhd	130
D. MHD Dynamos	131
1. cowling's theorem	131
2. parker's solar dynamo	132
E. Kinematic Dynamos	133
F. Mean-Field Dynamos	133
G. Astrophysical dynamos in the lab	134